

# Fast quantification of uncertainty in non-linear diffusion MRI models for artifact detection and more power in group studies

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## Abstract

Diffusion MRI (dMRI) allows for non-invasive investigation of brain tissue microstructure. By fitting a model to the dMRI signal, various quantitative measures can be derived from the data, such as fractional anisotropy, neurite density and axonal radii maps. The uncertainty in these dMRI measures is often ignored, while previous work in functional MRI has shown that incorporating uncertainty estimates can lead to group statistics with a higher statistical power. We propose the Fisher Information Matrix (FIM) as a generally applicable method for quantifying the parameter uncertainties in non-linear diffusion MRI models. In direct comparison with Markov Chain Monte Carlo sampling, the FIM produces similar uncertainty estimates at lower computational cost. Using acquired and simulated data, we then list several characteristics that influence the parameter variances, like data complexity and signal-to-noise ratio. In individual subjects, the parameter standard deviations can help in detecting white matter artifacts as patches of relatively large standard deviations. In group statistics, we recommend using the parameter standard deviations by means of variance weighted averaging. Doing so can reduce the overall variance in group statistics and reduce the effect of data artifacts without discarding data from the analysis. Both these effects can lead to a higher statistical power in group studies.

**Keywords:** Uncertainty estimates, Variances, Diffusion MRI, Microstructure, Fisher Information Matrix (FIM), Cramér Rao Lower Bound (CRLB)

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## <sup>1</sup> 1 Introduction

<sup>2</sup> Diffusion Magnetic Resonance Imaging (dMRI) allows for non-invasive in-  
<sup>3</sup> vestigation of brain tissue microstructure. By fitting a dMRI model to each

4 voxel, various quantitative measures can be derived from the data, such  
5 as fractional anisotropy (Basser et al., 1994), neurite density (Zhang et al.,  
6 2012) and axonal radii maps (Assaf & Pasternak, 2008; Alexander et al.,  
7 2010). These quantitative measures can be used in statistical group analysis.  
8 For example, tract-based spatial statistics (TBSS) is a popular approach to  
9 group analysis of fractional anisotropy measures (Smith et al., 2006). More  
10 often than not, these approaches (including TBSS) ignore the uncertainty in  
11 the quantitative measures. In functional magnetic resonance imaging, pre-  
12 vious work has shown that incorporating uncertainty estimates can lead to  
13 group statistics with a higher statistical power (Chen et al., 2012; Woolrich  
14 et al., 2004). For linear diffusion models, a method for computing and using  
15 uncertainty estimates has been shown before (Sjölund et al., 2018), but this  
16 has not yet been generalized to non-linear diffusion models like NODDI  
17 (Zhang et al., 2012) and CHARMED (Assaf & Basser, 2005).

18 Previous work in quantifying the parameter uncertainties include Markov  
19 Chain Monte Carlo (MCMC) (Behrens et al., 2003; Wegmann et al., 2017; Gu  
20 et al., 2017) and bootstrapping (Jones, 2003; Chung et al., 2006; Whitcher  
21 et al., 2008) methods. Of these two techniques, bootstrapping is often not  
22 applicable as it is either model specific (Whitcher et al., 2008) or requires  
23 very specific additional MRI measurements (Jones, 2003; Chung et al., 2006)  
24 which are often not acquired in diffusion MRI datasets. MCMC on the other  
25 hand can readily be extended to all microstructure models, but often re-  
26 quires long computation times, even with parallel processing on graphical  
27 processing units (Harms & Roebroeck, 2018).

28 We propose the Fisher Information Matrix (FIM) as a generally applicable  
29 method for quantifying the parameter uncertainties in non-linear diffusion  
30 MRI models. The FIM allows for estimating the local variances around the  
31 maximum likelihood point estimate, which is the point estimate typically  
32 used in group statistics. Computing the FIM is a relatively fast operation,  
33 requiring only a few additional model evaluations. In other fields, like for  
34 example astrophysics, the Fisher Information Matrix is already recognized  
35 as a useful tool for quantifying the uncertainty in parameter estimates (Val-  
36 lisneri, 2008; Rodriguez et al., 2013). In diffusion MRI, the FIM has been  
37 applied before, but only specific to the multi-Tensor model (Versteeg et al.,  
38 2018) and has not yet been generalized to all non-linear microstructure  
39 models.

40 The Fisher Information Matrix can additionally be used to compute the  
41 Cramér Rao Lower Bound (CRLB; Rao, 1945; Cramer, 1946), if and only  
42 if the true parameters are known (Kay, 1993). For example, in simula-  
43 tion studies the CRLB can function as a ground truth lower bound on  
44 the estimable variances, thereby indirectly evaluating the performance of  
45 the maximum likelihood routines (Kay, 1993). Although in brain data the

46 FIM can be interpreted as an approximation to the CRLB, we follow the re-  
47 sults in astrophysics and only interpret the FIM as a measure of uncertainty  
48 around the estimated parameters (Vallisneri, 2008).  
49 We first compare the uncertainty estimates from the Fisher Information  
50 Matrix to those of MCMC, using multiple datasets and multiple dMRI mi-  
51 crostructure models. We then investigate several data and model character-  
52 istics that can influence the parameter variances, like data complexity and  
53 Signal-to-Noise Ratio (SNR). In the end, we discuss the use of uncertainty  
54 estimates in white matter artifact detection (e.g. detecting fat saturation)  
55 and show how weighted averaging could lead to an increase in power in  
56 group studies.

## 57 2 Methods

### 58 2.1 Parameter distribution estimates

59 We compare two different methods for summarizing the parameter poste-  
60 rior distributions of a single voxel, a frequentist method using Maximum  
61 Likelihood Estimation (MLE) and the Fisher Information Matrix (FIM) and  
62 a Bayesian method using Markov Chain Monte Carlo (MCMC) (see fig-  
63 ure 1 for a schematic overview). With both methods we summarize the  
64 voxel-wise posteriors as a point estimate with a corresponding standard  
65 deviation (std.).

66 In the first method we use the Powell optimization routine (Powell, 1964;  
67 Harms et al., 2017) to get an MLE parameter point estimates. We estimate  
68 the standard deviations around those point estimates using the theory of  
69 the FIM. Standard deviations in derived parameter maps (e.g. Tensor Frac-  
70 tional Anisotropy) can be obtained by propagating the uncertainty of the  
71 model parameters. We refer to this method as MLE+FIM.

72 The second methodology uses MCMC sampling to approximate the full  
73 posterior distribution, using the Adaptive Metropolis Within Gibbs routine  
74 as discussed in (Harms & Roebroeck, 2018). From these samples we sum-  
75 marize the posterior distribution using a mean and standard deviation, as  
76 done before in before in dMRI modeling (Behrens et al., 2003; Sotiro-  
77 los et al., 2013; Wegmann et al., 2017). Uncertainties in derived parameter  
78 maps can be obtained by computing the derived parameter maps at ev-  
79 ery sampled point and summarizing the result. We refer to this method as  
80 MCMC.

81 The MLE+FIM provides a local variance around a mode while MCMC pro-  
82 vides a global variance around the mean. As such, these methods are only  
83 comparable if the posterior is unimodally Gaussian distributed, since then

84 the mean equals the mode. As in previous work (Behrens et al., 2003;  
85 Sotiropoulos et al., 2013; Wegmann et al., 2017), we assume the posteriors  
86 to be unimodally Gaussian distributed.

87 This assumption may not necessarily hold. For example, multi-modality  
88 could arise when fitting a single fiber model to a crossing fiber voxel. In  
89 such cases, different post-processing would be required on the MCMC sam-  
90 ples to correctly reflect the parameter variances. The FIM would be less  
91 sensitive to this issue since the FIM provides only local variances estimates.  
92 That is, the MLE would choose one mode of the distribution and the FIM  
93 would provide a local variance estimate around the chosen mode. This is-  
94 sue could also be circumvented by applying appropriate model selection to  
95 every voxel.

96 Non-Gaussian distributions can happen near parameter boundaries. For  
97 instance, very low (close to zero) or very high (close to one) compartment  
98 volume fractions can lead to a truncated posterior. In such cases the FIM  
99 no longer applies. For MCMC different post-processing would be required,  
100 like fitting a truncated normal distribution to the posterior. This could  
101 again be solved by appropriate model selection. We take no special pre-  
102 cautions for these boundary effects and assume these to not be present in  
103 white matter.

104 Nevertheless, we expect most posteriors to be unimodally Gaussian dis-  
105 tributed. This assumption is also supported by two theoretical arguments.  
106 First, if the model is suitable to describe the data (e.g. if model selection was  
107 successfully applied), the posterior asymptotically approaches a Gaussian  
108 distribution (Gelman et al., 2013). Second, according to the central limit  
109 theorem, each parameter's marginal distribution will asymptotically tend  
110 to a Gaussian as the number of model parameters increases (Gelman et al.,  
111 2013).

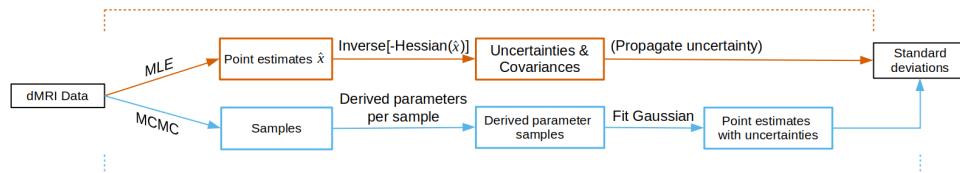


Figure 1: The uncertainty computation methods for both the Maximum Likelihood Estimation (MLE) and Markov Chain Monte Carlo (MCMC) methods.

### 112 2.1.1 Fisher Information Matrix

113 The observed Fisher Information Matrix is defined as the negative Hessian  
114 of the log-likelihood function when evaluated at the maximum likelihood  
115 estimate (Pawitan, 2013; Gelman et al., 2013). The inverse of the FIM is

116 an asymptotic estimator of the covariance matrix (Pawitan, 2013; Gelman  
 117 et al., 2013). Formally, let  $l(\mathbf{x})$  be a log-likelihood function with maximum  
 118 likelihood estimate  $\hat{\mathbf{x}}$ . A second order Taylor approximation of  $l(\mathbf{x})$  cen-  
 119 tered at  $\hat{\mathbf{x}}$  is then given by:

$$l(\mathbf{x}) = l(\hat{\mathbf{x}}) + \frac{1}{2}(\mathbf{x} - \hat{\mathbf{x}})^T \frac{\partial^2}{\partial \mathbf{x}^2} l(\hat{\mathbf{x}})(\mathbf{x} - \hat{\mathbf{x}}) \quad (1)$$

120 ignoring the higher terms and having dropped the linear term since the  
 121 first derivative of a function is zero at the mode. Considering the first term,  
 122  $l(\hat{\mathbf{x}})$ , a constant and the second term,  $\frac{1}{2}(\mathbf{x} - \hat{\mathbf{x}})^T \frac{\partial^2}{\partial \mathbf{x}^2} l(\hat{\mathbf{x}})(\mathbf{x} - \hat{\mathbf{x}})$ , proportional  
 123 to the logarithm of a normal density, we get the approximation:

$$l(\mathbf{x}) \approx \mathcal{N}(\hat{\mathbf{x}}, [\mathbf{I}(\hat{\mathbf{x}})]^{-1}) \quad (2)$$

124 where  $\mathbf{I}(\hat{\mathbf{x}})$  is the observed Fisher Information Matrix:

$$\mathbf{I}(\hat{\mathbf{x}}) = -\mathbf{H}(\hat{\mathbf{x}}) = -\frac{\partial^2}{\partial \mathbf{x}^2} l(\hat{\mathbf{x}}) \quad (3)$$

125 For the Hessian to be positive definite, this theory requires  $\hat{\mathbf{x}}$  to lie within  
 126 the boundaries of the parameter space (Gelman et al., 2013). We compute  
 127 the Hessian numerically (see Appendix A) and its inverse using a direct  
 128 inverse where possible with a fallback on the (Moore-Penrose) pseudo-  
 129 inverse for ill-conditioned Hessians. Ill-conditioned Hessian can for exam-  
 130 ple arise with parameter estimates lying at a predefined parameter bound-  
 131 ary (Gelman et al., 2013).

### 132 2.1.2 *Uncertainty propagation*

133 Given a function  $\mathbf{y} = f(\boldsymbol{\theta})$  where  $f(\cdot)$  is a known function, uncertainty  
 134 propagation provides the probability distribution of  $\mathbf{y}$  given the probability  
 135 distribution of  $\boldsymbol{\theta}$ . For example, we can use this to estimate the standard  
 136 deviation of a Tensor Fractional Anisotropy (FA) estimate, by propagating  
 137 the standard deviation estimates of the Tensor diffusivities. We use a first  
 138 order Taylor expansion linear approximation (Arras, 1998), which states  
 139 that if  $\boldsymbol{\theta}$  is normally distributed with mean  $\boldsymbol{\mu}_{\boldsymbol{\theta}}$  and covariance matrix  $\Sigma_{\boldsymbol{\theta}}$ ,  
 140 the distribution of  $\mathbf{y}$  can be approximated as:

$$\mathbf{y} \approx \mathcal{N}(\boldsymbol{\mu}_{\mathbf{y}}, \Sigma_{\mathbf{y}}) = \mathcal{N}(f(\boldsymbol{\theta}), \mathbf{J}_f(\boldsymbol{\theta}) \Sigma_{\boldsymbol{\theta}} \mathbf{J}_f(\boldsymbol{\theta})^T) \quad (4)$$

<sup>141</sup> with  $\mathbf{J}_f$  the Jacobian matrix of  $f$ . More succinctly, the covariance matrix of  
<sup>142</sup>  $\mathbf{y} = f(\boldsymbol{\theta})$  is given by:

$$\Sigma_{\mathbf{y}} = \mathbf{J}_f \Sigma_{\boldsymbol{\theta}} \mathbf{J}_f^\top \quad (5)$$

<sup>143</sup> which holds as a generally applicable formula for linear propagation of co-  
<sup>144</sup> variances (Arras, 1998). In the case of an univariate output  $y = f(\boldsymbol{\theta})$ , the  
<sup>145</sup> Jacobian can be formulated as a gradient vector  $\nabla_f$ , leading to the follow-  
<sup>146</sup> ing expression for the variance in  $y$ :

$$\sigma_y^2 = \nabla_f \Sigma_{\boldsymbol{\theta}} \nabla_f^\top \quad (6)$$

<sup>147</sup> This error propagation technique uses both the variances and the co-variances  
<sup>148</sup> of all the propagated parameters. Additionally, this technique takes into ac-  
<sup>149</sup> count the functional form of the propagated function, i.e. if the function is  
<sup>150</sup> linear or non-linear. The Jacobian or gradient can be computed numerically  
<sup>151</sup> using finite-differences or can be evaluated at an analytical derivative. We  
<sup>152</sup> use analytical expressions for all uncertainty propagations. See Appendix B  
<sup>153</sup> for worked out error propagation examples of the Tensor FA and Ball&Stick  
<sup>154</sup> Fraction of Stick parameters.

## <sup>155</sup> 2.2 Variance weighted average

<sup>156</sup> Variance weighted averaging makes it possible to include the variances of  
<sup>157</sup> the data points when computing a mean and standard deviation. For ex-  
<sup>158</sup> ample, the voxel-wise variances discussed earlier can be used in averages  
<sup>159</sup> of white matter regions within a subject, or in voxel-wise averages over  
<sup>160</sup> multiple subjects. First, given  $n$  data points  $z_i$ , we define the regular mean  
<sup>161</sup> as:

$$\bar{\mu}_{\text{regular}} = \frac{1}{n} \sum_i^n z_i \quad (7)$$

<sup>162</sup> and regular standard deviation as:

$$\bar{\sigma}_{\text{regular}} = \sqrt{\frac{\sum_i^n (z_i - \bar{\mu}_{\text{regular}})^2}{n}} \quad (8)$$

<sup>163</sup> If each data point  $z_i$  has a corresponding weight  $w_i$ , we can compute a  
<sup>164</sup> weighted mean as:

$$\bar{\mu}_{\text{weighted}} = \frac{\sum_i^n w_i z_i}{\sum_i^n w_i} \quad (9)$$

165 and a weighted standard deviation as:

$$\bar{\sigma}_{\text{weighted}} = \sqrt{\frac{\sum_i^n w_i(z_i - \bar{\mu}_{\text{weighted}})^2}{\frac{(m-1)}{m} \sum_i^n w_i}} \quad (10)$$

166 with  $m$  for the number of non-zero weights, included here to allow for non-  
167 normalized weights. It has been shown that the weights that minimize the  
168 variance of the weighted average are the reciprocals of the variances of each  
169 of the data points  $z_i$  (Shahar, 2017). That is, given the variances  $\sigma_i^2$  for each  
170  $z_i$ , the weights that minimize  $\text{Var}(\sum_i w_i z_i)$  is given by:

$$w_i = \frac{1}{\sigma_i^2} \quad (11)$$

171 Incidentally, these weights are also the maximum likelihood estimator of  
172 the weighted mean and variance under the assumption that the data points  
173  $z_i$  are independent and normally distributed with the same mean (Cochran,  
174 1937).

### 175 2.3 *Diffusion microstructure models*

176 To capture the variety of microstructure models in diffusion MRI we chose  
177 four different models, the Tensor (Basser et al., 1994), Ball&Stick (Behrens  
178 et al., 2003), Bingham-NODDI (Tariq et al., 2016) and CHARMED (Assaf  
179 et al., 2004) models. The Tensor model is the oldest diffusion MRI model  
180 and still sees widespread usage in the literature. From the Tensor we derive  
181 the Fractional Anisotropy (FA) quantity. The Ball&Stick model (Behrens  
182 et al., 2003) is the first multi-compartment model and is often used as lo-  
183 cal estimator for tractography. To delineate multiple fiber orientations,  
184 the Ball&Stick model can feature multiple Stick compartments, but always  
185 with a single Ball compartment. To differentiate between the Ball&Stick  
186 models with one or more Stick compartments, we denote the specific Ball&Stick  
187 model as "BallStick\_in1", "BallStick\_in2" and "BallStick\_in3" for respec-  
188 tively one, two or three Stick compartments. This is a general naming  
189 scheme to denote models that can have one or more intra neuronal com-  
190 partments relative to the other compartments. From the Ball&Stick model  
191 we derive the Fraction of Stick (FS) quantity, which is the sum of the vol-  
192 ume fractions of the Stick compartments.

193 More recent, biologically inspired, models include Bingham-NODDI and  
194 CHARMED. The Bingham-NODDI model assumes that white matter con-  
195 sists of restricted intra-cellular and hindered extra-cellular water compart-  
196 ments, with the intra-cellular compartment capturing neurite orientation

197 dispersion. From the Bingham-NODDI model we use the Fraction of Re-  
198 stricted (FR) quantity, the volume fraction of the restricted intra-cellular  
199 compartment. The CHARMED model assumes a tissue model of restricted  
200 intra-neuronal and hindered extra-neuronal water compartments, with the  
201 intra-neuronal compartment assuming a bundle of axons. Since CHARMED  
202 can be used with multiple intra-neuronal compartments we again denote  
203 these with the 'in' suffix. Here, we only use CHARMED with one intra-  
204 neuronal compartment, denoted as "CHARMED\_in1". From the CHARMED  
205 model we use the Fraction of Restricted (FR) quantity, the volume fraction  
206 of the restricted intra-neuronal compartment. For implementation notes of  
207 these models see (Harms et al., 2017).

208 *2.4 Software*

209 All models and routines used in this study are implemented in a Python  
210 based GPU (graphical processing unit) accelerated toolbox, the Microstruc-  
211 ture Diffusion Toolbox, MDT, which is freely available under an open source  
212 license at <https://github.com/cbclab/MDT>. We used the models and  
213 MCMC routine as implemented in MDT version 0.18.3. From this version  
214 onward, MDT automatically computes the FIM after every maximum like-  
215 lihood estimation operation and writes out the variances and covariances  
216 alongside the parameter estimates. Scripts for reproducing the results in  
217 this article can be found at [https://github.com/robbert-harms/uncertainty\\_paper](https://github.com/robbert-harms/uncertainty_paper). All computations for this paper were performed  
219 on a single AMD Fury X graphics card.

220 *2.5 Datasets*

221 In this study we used simulated data and imaging data from two popula-  
222 tion studies. To illustrate the methods on a dataset with a clinically feasible,  
223 fast to acquire, acquisition scheme, we used data from the diffusion pro-  
224 tocol pilot phase of the Rhineland Study ([www.rheinland-studie.de](http://www.rheinland-studie.de)).  
225 We refer to these datasets and acquisition schemes as *RLS-pilot*. To illustrate  
226 the methods on a dataset with a high-end, long acquisition time, acquisition  
227 scheme, we used data from the Human Connectome Project MGH-USC  
228 Young Adult study. We refer to these datasets and acquisition schemes as  
229 *HCP MGH*. For simulated data we used a single representative acquisition  
230 scheme from both the RLS-pilot and HCP MGH studies.

231 The RLS-pilot datasets were acquired on a Siemens Magnetom Prisma (Siemens,  
232 Erlangen, Germany) with the Center for Magnetic Resonance Research (CMRR)  
233 multi-band (MB) diffusion sequence (Moeller et al., 2010; Xu et al., 2013).  
234 These datasets had a resolution of 2.0 mm isotropic with diffusion param-  
235 eters  $\Delta = 45.8$  ms,  $\delta = 16.3$  ms, TE = 90 ms and TR = 4500 ms, and with

236 Partial Fourier = 6/8, MB factor 3, no in-plane acceleration with 3 shells of  
237  $b = 1000, 2000, 3000 \text{ s/mm}^2$ , with respectively 30, 40 and 50 directions to  
238 which are added 14 interleaved b0 volumes leading to 134 volumes in total  
239 per subject. Additional b0 volumes were acquired with a reversed phase  
240 encoding direction which were used to correct susceptibility related distor-  
241 tion (in addition to bulk subject motion) with the topup and eddy tools in  
242 FSL version 5.0.9 (Andersson & Sotiroopoulos, 2016). The total acquisition  
243 time is 10 min 21 sec. These three-shell datasets represent a relatively short  
244 time acquisition protocol that still allows many models to be fitted. From  
245 this dataset we used a single representative subject (v3a\_1.data\_ms20).

246 The HCP MGH datasets come from the freely available fully preprocessed  
247 dMRI data from the USC-Harvard consortium of the Human Connectome  
248 project. Data used in the preparation of this work were obtained from the  
249 MGH-USC Human Connectome Project (HCP) database (<https://ida.loni.usc.edu/login.jsp>). The data were acquired on a specialized  
250 Siemens Magnetom Connectom with 300 mT/m gradient set (Siemens, Er-  
251 langen, Germany). These datasets were acquired at a resolution of 1.5 mm  
252 isotropic with diffusion parameters  $\Delta = 21.8 \text{ ms}$ ,  $\delta = 12.9 \text{ ms}$ ,  $TE = 57 \text{ ms}$ ,  
253  $TR = 8800 \text{ ms}$ , Partial Fourier = 6/8, MB factor 1 (i.e. no simultaneous  
254 multi-slice), in-plane GRAPPA acceleration factor 3, with 4 shells of  $b =$   
255  $1000, 3000, 5000, 10,000 \text{ s/mm}^2$ , with respectively 64, 64, 128, 393 directions  
256 to which are added 40 interleaved b0 volumes leading to 552 volumes in  
257 total per subject, with an acquisition time of 89 minutes. These four-shell,  
258 high number of directions, and very high maximum b-value datasets allow  
259 a wide range of models to be fitted. From these datasets we used a single  
260 representative subject (hcp\_1003) in single subject illustrations and we used  
261 all 35 subjects in the group comparisons.

262 Since the CHARMED\_in1 model requires relatively high b-values ( $\geq \sim 6000$   
263  $\text{s/mm}^2$ ), which are not present in the RLS-pilot datasets, we will only use  
264 the HCP MGH dataset when showing CHARMED\_in1 results. Addition-  
265 ally, since the Tensor model is only valid for b-values up to about 1200  
266  $\text{s/mm}^2$ , we only use the b-value  $1000 \text{ s/mm}^2$  shell and b0 volumes in maxi-  
267 mum likelihood estimation and MCMC sampling of the Tensor model. All  
268 other models use all the data volumes.

269 For all datasets we created a white matter (WM) mask from the Tensor FA  
270 estimates and, using BET from FSL (Smith, 2002), a whole brain mask. The  
271 whole brain mask is used for MLE and MCMC sampling, whereas aver-  
272 ages over the WM mask are used in model or data comparisons. For each  
273 dataset, voxel-wise SNR is estimated using only the unweighted (b0) vol-  
274 umes, by dividing the mean of the unweighted volumes by the standard  
275 deviation.

277 2.5.1 *Ground truth simulations*

278 We additionally created simulated data to illustrate the effects of the signal-  
279 to-noise ratio (SNR) on the variance of the estimated parameters. We used a  
280 single representative acquisition scheme from both the RLS-pilot and HCP  
281 MGH datasets (the acquisition schemes of subject v3a\_1\_data\_ms20 and  
282 hcp\_1003), and simulated data for each model. For each acquisition scheme  
283 and each model, we simulate 10000 voxels with random volume fractions  
284 in  $[0.2, 0.8]$ , diffusivities in  $[5e-11, 5e-9]$  mm $^2$ /s, and orientations in  $[0, \pi]$ .  
285 From these, we created multiple copies with Rician noise (Gudbjartsson &  
286 Patz, 1995) of SNRs 5, 10, 20, 30, 40 and 50. We then fit and sample each  
287 model ten times to these simulated datasets and estimate the standard de-  
288 viation using both the FIM and MCMC approach as described above. Per  
289 SNR we summarize the results of these ten trials as a mean standard devi-  
290 ation and its corresponding standard error of the mean.

291 2.5.2 *Group statistics*

292 For the group statistics we computed Tensor FA and Bingham-NODDI FR  
293 and FR standard deviation maps on all 35 subjects using the MLE+FIM  
294 method. To be able to compare the subjects, we first registered the Tensor  
295 FA estimates to the FMRIB58\_FA\_1mm template using FLIRT and FNIRT  
296 from FSL (Andersson et al., 2010). Next, we used those registration tem-  
297 plates to co-register the Bingham-NODDI FR and FR standard deviation  
298 maps.

299 With uncertainty maps available there are three methods to compute group  
300 statistics that are robust against subject-level artifacts. Method one, ap-  
301 ply variance weighted averaging using the uncertainty estimates to down-  
302 weight voxels with a high standard deviation. This would automatically re-  
303 move artifacts if these artifacts lead to high parameter uncertainties. Method  
304 two, exclude outlier subjects from the group statistic. Outlier subjects could  
305 be detected using the point estimates or using the uncertainty maps. Method  
306 three, use a combination of method one and two, i.e. computing weighted  
307 group estimates after removal of outliers.

308 To illustrate these three artifact reduction methods, we first computed a  
309 baseline statistic using a simple mean and standard deviation over all 35  
310 subjects. We then used artifact reduction method one and used the FR stan-  
311 dard deviation maps as weights in the variance weighted averaging. To  
312 apply artifact reduction method two and three, we created a new subgroup  
313 with only 30 subjects, where we manually removed five subjects (mgh\_-  
314 1008, mgh\_1009, mgh\_1013, mgh\_1017 and mgh\_1032) that had a large white  
315 matter artifact over the corpus callosum. We then applied regular averag-  
316 ing and weighted averaging over these remaining 30 subjects.

317 As a comparison method between regular and weighted averaging we com-  
318 puted  $(\mu_{\text{weighted}} - \mu_{\text{regular}})/\mu_{\text{regular}}$  and  $(\sigma_{\text{weighted}} - \sigma_{\text{regular}})/\sigma_{\text{regular}}$  as dif-  
319 ference measure for the mean and standard deviation estimates between  
320 regular and weighted averaging.

### 321 3 Results

322 We begin by comparing the parameter estimates and parameter uncertainty  
323 estimates of MLE+FIM to the corresponding estimates from MCMC. Next,  
324 we investigate the effect of SNR on the parameter standard deviations us-  
325 ing both simulated and imaging data. We end with a comparison of regular  
326 versus weighted averaging in group statistics.

#### 327 3.1 Parameter distribution estimates

328 Figure 2 visually compares the results of MLE+FIM to those of MCMC, us-  
329 ing the Bingham-NODDI Fraction of Restricted (FR) parameter, on a single  
330 subject from the RLS-pilot dataset. Comparing results of a single transverse  
331 slice shows high qualitative correspondence between the MLE and MCMC  
332 methods (figure 2A), with both the point estimates and corresponding stan-  
333 dard deviations (stds.) in close resemblance. A single voxel illustration of  
334 the estimated Gaussian distributions (figure 2B) again shows a high degree  
335 of similarity, with both Gaussian fits capturing the characteristics of the  
336 MCMC sample distribution to a large degree.

337 To further quantify the correspondence between the MLE and MCMC method-  
338 ologies, we created scatter plots between the MLE and MCMC estimates of  
339 both the point estimate and standard deviation estimate. This was per-  
340 formed over a white matter mask for a single subject from both the HCP  
341 MGH and RLS-pilot datasets. Figure 3 shows Bingham-NODDI FR mean  
342 and standard deviation scatter plots. The FR point estimates are very tightly  
343 confined to the identity line, illustrating a high degree of correspondence in  
344 the point estimates from MCMC and MLE. The variation of point estimates  
345 along the diagonal corresponds to variation of FR values over the white  
346 matter mask, ranging between roughly 0.3 and 0.7. The std. estimates be-  
347 tween the MLE and MCMC methodologies again show a high correspon-  
348 dence, although the off-diagonal spread in the std. plot is visibly larger  
349 than that in the point estimate plot. There is also some clipping visible in  
350 the std. plot, with MLE estimating a zero std. while MCMC provides a  
351 range of values. This is mostly due to very low point estimates (near zero),  
352 at which point the FIM is no longer applicable. The blue-green-yellow-red  
353 coded points in both plots account for 97-99.5% of the voxels and the pur-  
354 ple points account for the remaining fraction of outliers. The std. estimates

355 for the HCP MGH data are clearly lower than for the RLS-pilot data, con-  
356 firming an expected higher precision (lower uncertainty) of point estimates  
357 based on more dMRI data-points.

358 To investigate the correspondence in MCMC and MLE uncertainty esti-  
359 mates for a larger number of models, figure 4 shows scatter plots for mul-  
360 tiple microstructure models. Parameter point estimate comparisons are  
361 not shown here, but are generally in correspondence to a very high de-  
362 gree. Across all models and data, except for the CHARMED\_in1 model  
363 fit on RLS-pilot data, MCMC and MLE uncertainty estimates are in high  
364 correspondence and located close to the identity diagonal. A relatively  
365 large off-diagonal variance in standard deviation estimates is visible in the  
366 CHARMED\_in1 FR parameter on the RLS-pilot data. This is expected as the  
367 RLS-pilot dataset is not well suited for the CHARMED\_in1 model due to  
368 too low b-values (the CHARMED\_in1 model requires  $b\text{-values} \leq 6000\text{s/mm}^2$ ).  
369 Standard deviation estimates for CHARMED\_in1 on the HCP MGH data  
370 are not only much more tightly confined to the identity diagonal, the std.  
371 estimates themselves are also about a factor two lower. A large spread to  
372 the right is also visible in the Ball&Stick\_in3 results. This might be related  
373 to MLE choosing a different mode and is perhaps solved using model se-  
374 lection. There is also again some clipping visible, with MLE providing a  
375 zero std. with voxels with a very low point estimate.

376 Irrespective of the method (MCMC or MLE+FIM), the std. estimates on  
377 the RLS-pilot data are always higher than the corresponding estimates on  
378 the HCP MGH data, once again confirming the expected higher precision  
379 on datasets with a larger number of direction. Conversely, one would ex-  
380 pect higher complexity models (i.e. models with more compartments and  
381 more parameters to fit) to have higher uncertainty when fitted on the same  
382 data. This is indeed illustrated by the Ball&Stick\_in{1,2,3} results, were we  
383 see an increasing estimated standard deviation for an increasing number  
384 of Sticks, within each of the HCP MGH and RLS-pilot datasets. Finally,  
385 Tensor FA standard deviations are about a factor two higher than those of  
386 the other models. This is probably related to Tensor FA being a compound  
387 parameter.

388 To quantify correspondence in the MCMC and MLE std. estimates in the  
389 scatter plots, table 1 shows the percentage of voxels for which the differ-  
390 ence between the MLE and MCMC variances is less than two standard de-  
391 viations from the mean difference. We note an average similarity of  $\sim 98.7\%$   
392 across six models and two datasets, even including the 97.9% similarity for  
393 the CHARMED\_in1 model fit on RLS-pilot data. Table 2 compares runtimes  
394 between the MLE with the FIM and the MCMC methodologies, measuring  
395 the time between loading the data and writing the results. Averaged over  
396 six models and two subjects, the GPU-optimized MLE and FIM together

397 compute approximately 38 times faster than GPU-optimized MCMC.

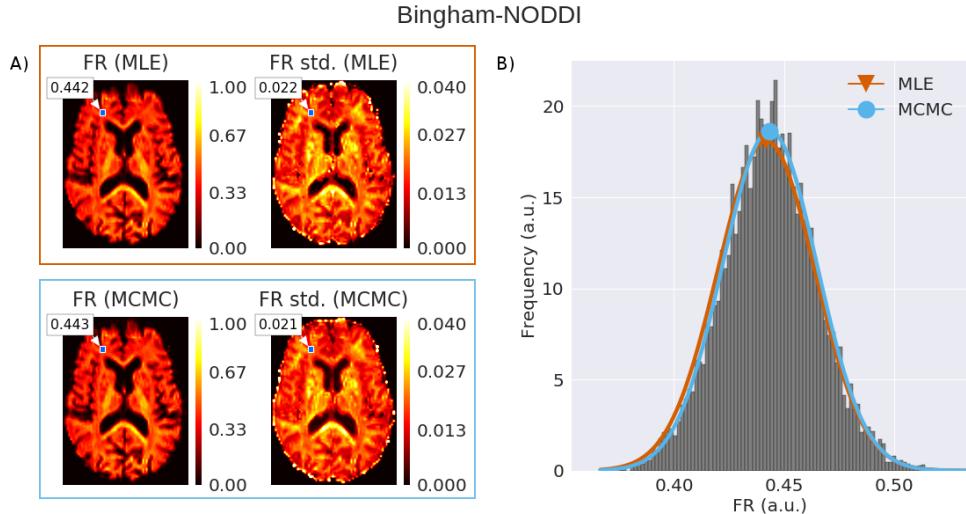


Figure 2: A) Visual comparison of parameter and standard deviation uncertainty maps between the Maximum Likelihood Estimation (MLE) and Markov Chain Monte Carlo (MCMC) methodologies for the Bingham-NODDI Fraction of Restricted (FR) on an RLS-pilot dataset. B) Histogram of the 20 thousand MCMC samples of the highlighted voxel in figure A, with in red and blue the fitted Gaussian distributions of, respectively, the MLE and MCMC methodologies.

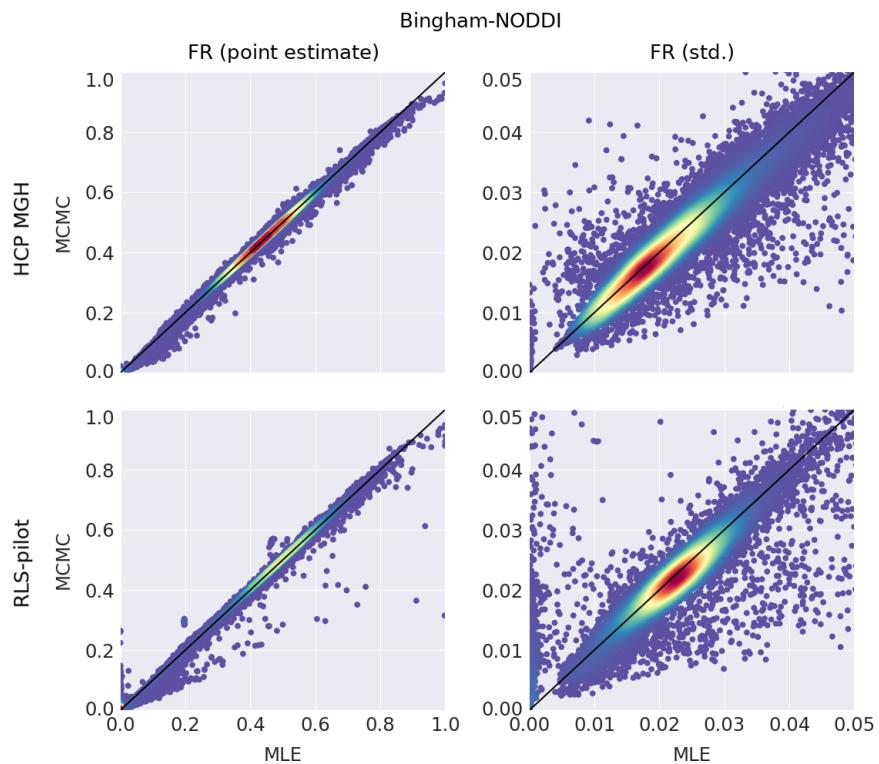


Figure 3: Scatter plots comparing Maximum Likelihood Estimation (MLE) and Markov Chain Monte Carlo (MCMC) point estimates (left column) and standard deviations (right column) for the Bingham-NODDI Fraction of Restricted (FR) values over a white matter mask for both a complex, long acquisition time HCP MGH dataset and a clinically feasible RLS-pilot dataset. Plots are color coded using a kernel density estimate (a.u) from purple (low density) to red (high density). Purple points correspond to a small percentage (0.5-3%) of the data (c.f. Table 1).

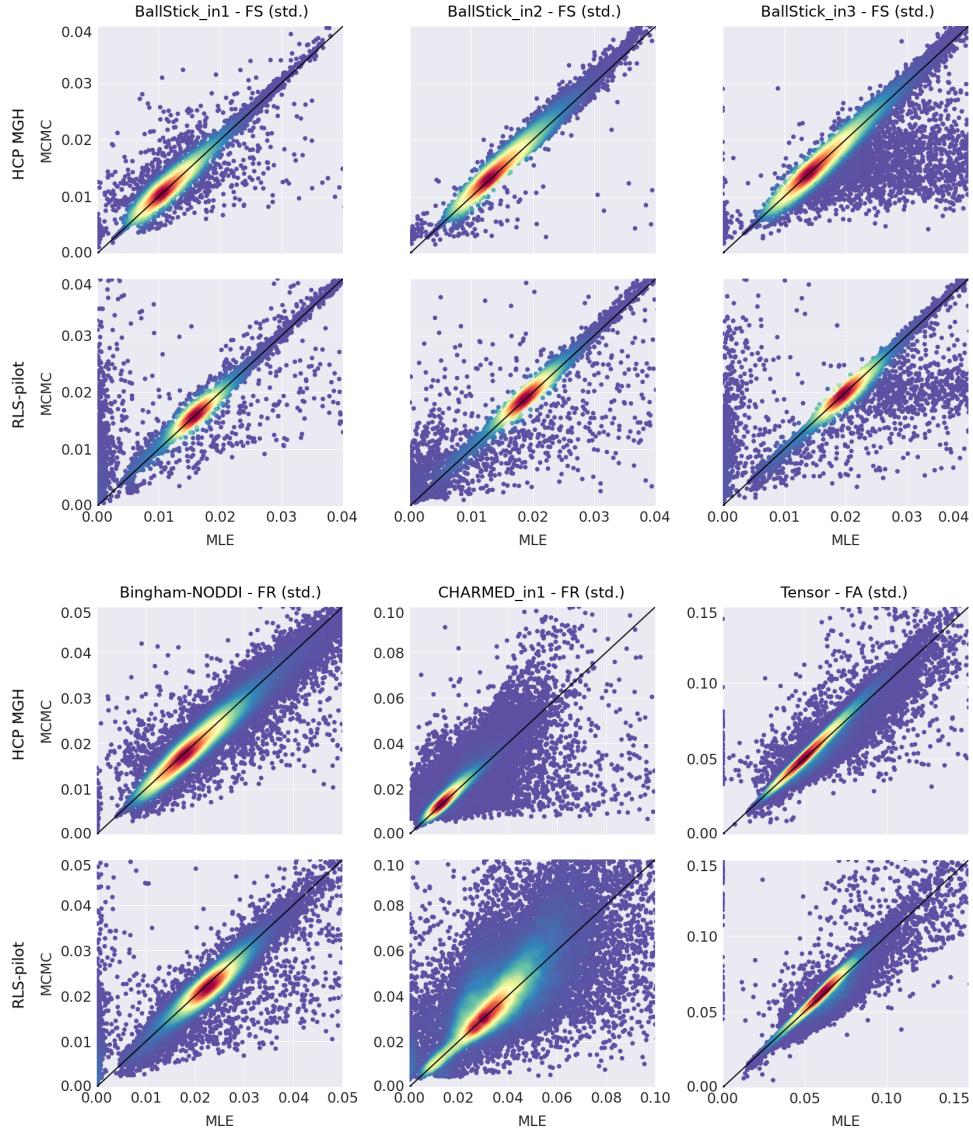


Figure 4: Scatter plots comparing Maximum Likelihood Estimation (MLE) and Markov Chain Monte Carlo (MCMC) standard deviations for multiple models over a white matter mask for both an HCP MGH and an RLS-pilot dataset. Acronyms are Fraction of Stick (FS), Fraction of Restricted (FR) and Fractional Anisotropy (FA). Plots are color coded using a kernel density estimate (a.u) from purple (low density) to red (high density). Purple points correspond to a small percentage (0.5-3%) of the data (c.f. Table 1).

	HCP MGH	RLS-pilot
Ball&Stick_in1	99.5%	98.8%
Ball&Stick_in2	99.9%	99.4%
Ball&Stick_in3	98.8%	97.5%
Bingham-NODDI	98.9%	98.7%
CHARMED_in1	98.6%	96.9%
Tensor	99.0%	97.9%

Table 1: For each model and dataset the percentage of voxels where the difference between the parameter stds. from the FIM and of MCMC are within two standard deviations from the mean difference. These percentages correspond to the red/yellow high densities in figure 4.

	HCP MGH		RLS-pilot			
	MLE + FIM	MCMC	rel.	MLE + FIM	MCMC	rel.
Ball&Stick_in1	00:01:49	01:21:55	45x	00:00:30	00:20:49	42x
Ball&Stick_in2	00:04:32	02:33:18	34x	00:01:08	00:42:36	38x
Ball&Stick_in3	00:13:01	07:00:51	32x	00:03:19	01:53:33	34x
Bingham-NODDI	02:06:19	111:32:52	53x	00:28:19	26:11:47	56x
CHARMED_in1	02:09:49	53:34:47	25x	00:21:53	07:49:55	21x
Tensor	00:02:41	01:59:07	44x	00:02:18	01:02:11	27x

Table 2: Runtime comparison between the two methodologies for computing parameter statistics, Maximum Likelihood Estimation (MLE) with the Fisher Information Matrix (FIM) and Markov Chain Monte Carlo (MCMC) sampling, for six different models and using a single representative subject from both the HCP MGH and the RLS-pilot datasets. Reported run times are over the entire brain mask and are in units of (h:m:s), with next to it the relative speed advantage of the MLE + FIM over MCMC.

398 3.2 *Effect of SNR on parameter variances*

399 Lower SNR per data point (i.e. single diffusion volume) is expected to lead  
400 to higher uncertainty in fitted parameter estimates. This issue is of extra im-  
401 portance in brain dMRI by the fact that SNR is non-uniform over the brain,  
402 especially in modern high number-of-channel phased array RF-coils. In or-  
403 der to assess the effect of SNR on parameter variances, figure 5 compares  
404 an estimate of SNR, its reciprocal, and the parameter standard deviation  
405 estimates of multiple white matter models on a single HCP MGH dataset.  
406 We observe a decreased SNR in the center of the brain and an increase of  
407 SNR towards the periphery. A very similar gradient can be observed in the  
408 standard deviation maps, with a decrease in parameter standard deviations  
409 towards the periphery. As in the previous results, we observe an increase  
410 in standard deviations for an increased number of Sticks in the Ball&Stick\_-  
411 in{1,2,3} models, and Tensor FA standard deviations are about a factor two  
412 higher than the other standard deviation estimates.

413 To further compare SNR and standard deviation estimates, figure 6 plots  
414 SNR versus parameter standard deviations, for both simulated data and  
415 imaging data. In general, we observe an inverse relationship between SNR  
416 and standard deviation, where an increase in SNR leads to an decrease  
417 in parameter std. estimates. Standard deviations on RLS-pilot data are  
418 always higher than corresponding estimates on HCP MGH data, except  
419 for the imaged data analysis at an SNR of 5, where the RLS-pilot dataset  
420 has a lower standard deviation. For lower SNR (< 10), MLE std. esti-  
421 mates are slightly higher than the MCMC estimates. For higher SNR (>  
422 10), the MLE and MCMC standard deviation estimates quickly converge,  
423 except for Ball&Stick\_in2, Ball&Stick\_in3 and Tensor estimates on the RLS-  
424 pilot dataset, where MLE standard deviations stay higher than those from  
425 MCMC. For the HCP MGH dataset, results are consistent between simu-  
426 lated and imaging data, with differences within the Standard Error of the  
427 Mean (SEM). Results on the RLS-pilot dataset are generally also consistent,  
428 except for an SNR of 5, where imaging data results are lower than those  
429 on simulated data. We finally observe that the standard error of the mean  
430 is generally higher for the simulated data compared to the imaging data,  
431 especially for lower SNR.

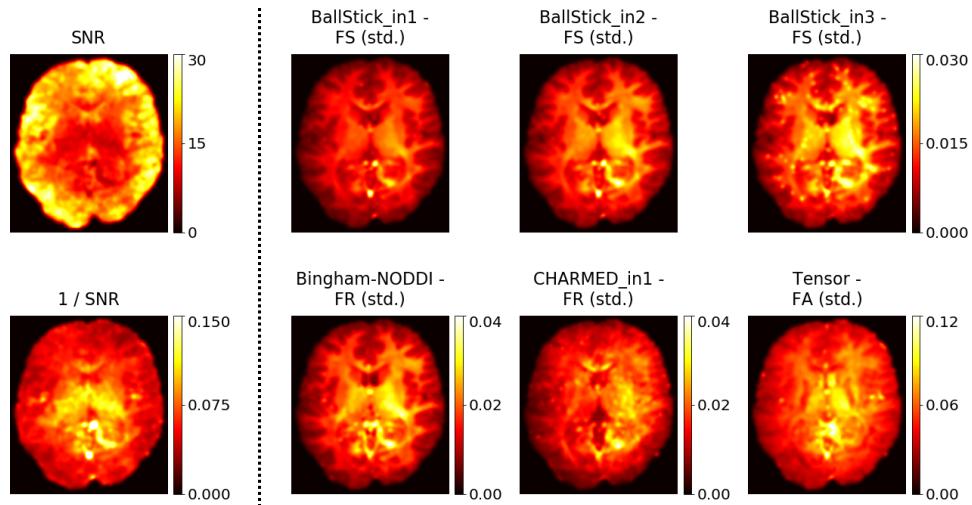


Figure 5: Illustration of the effect of Signal to Noise ratio (SNR) on parameter standard deviation estimates (using the MLE methodology), for a single HCP MGH subject (subject 1003). Maps are slightly smoothed with a 3d Gaussian filter ( $\sigma = 1\text{voxel}$ ). Parameter acronyms are Fraction of Stick (FS), Fraction of Restricted (FR) and Fractional Anisotropy (FA).

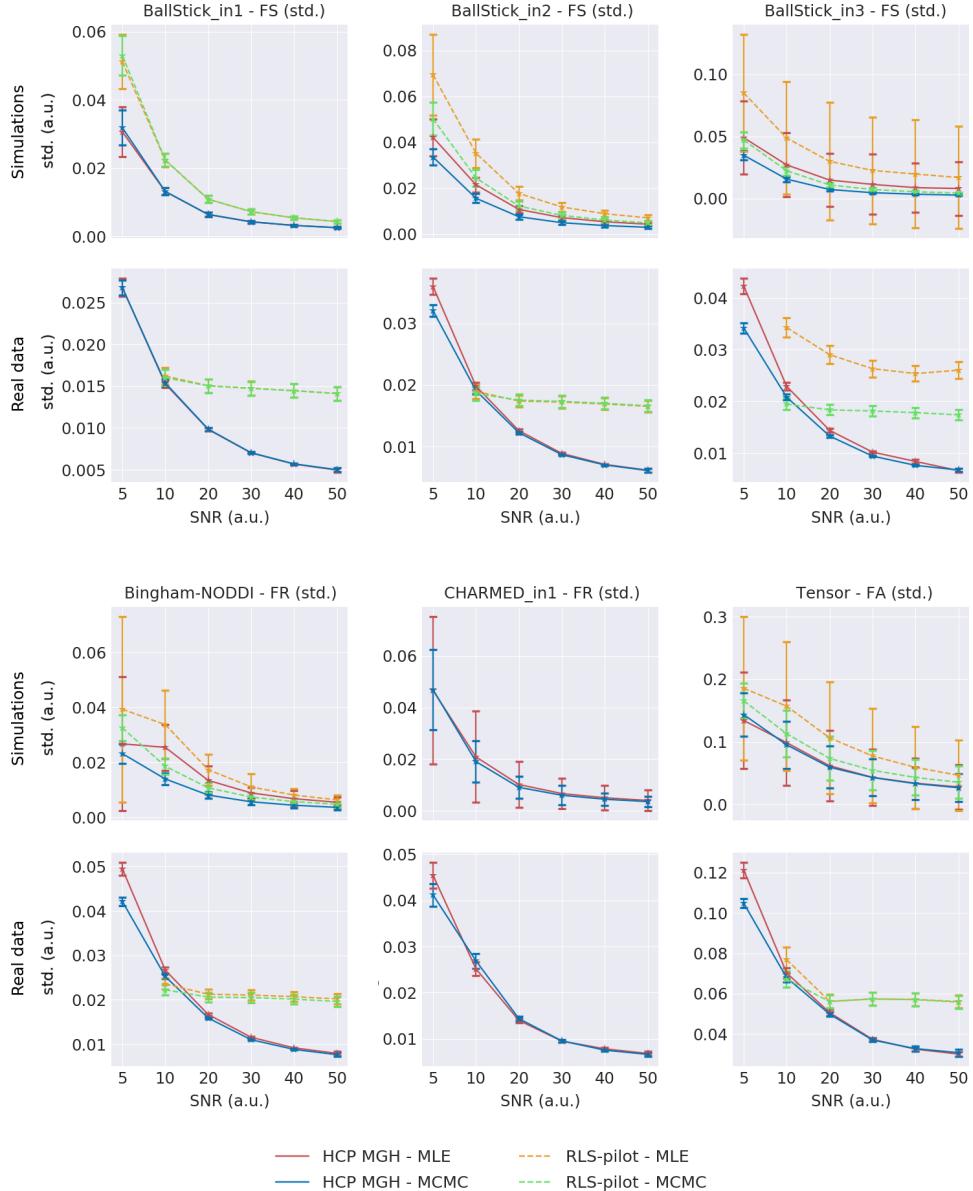


Figure 6: Effect of SNR on parameter standard deviations for simulated data and imaging data. Simulation results are over 10000 simulated voxels per SNR with a standard error of the mean (SEM) as error bar over 10 optimization and sampling trials. Real data results are for 10 subjects of the HCP MGH and 10 subjects of the RLS-pilot datasets, with SNR estimated as  $\text{mean}(\text{b0\_volumes})/\text{std}(\text{b0\_volumes})$ .

432 3.3 *Group statistics*

433 Figure 7 shows Bingham-NODDI FR results of three subjects of the HCP  
434 MGH dataset after co-registration, to illustrate the behavior of standard  
435 deviations in regions of white matter acquisition artifacts. The first sub-  
436 ject (top row) has a clear artifact across the genu of the corpus callosum,  
437 perhaps due to incomplete fat saturation. This artifact is visible in both  
438 the mean parameter estimates and the standard deviation estimates. The  
439 second subject (middle row) shows a patch of relatively large standard de-  
440 viations in and near the splenium of the corpus callosum, without an eas-  
441 ily detectable alteration in the mean parameter map. For comparison, we  
442 show a third subject (bottom row) at the same contrast scaling, with no  
443 visible artifacts or alterations in either the mean or standard deviation es-  
444 timates. This figure illustrates that parameter std. maps can play a role  
445 in detecting biased estimates resulting from imaging artifacts. In particu-  
446 lar, artifacts which may not always be detectable in the parameter maps  
447 themselves.

448 Figure 8 shows four group statistic estimates, a regular (baseline) and three  
449 statistics using the three mentioned artifact reduction methods using the  
450 parameter variances. To reiterate, these were method one, a weighted av-  
451 erage on all 35 subjects, method two, remove outlier subjects and apply  
452 regular averaging and method three, a weighted average with outlier sub-  
453 jects removed. Between regular and weighted averaging we computed a  
454 percentile difference map over a white matter mask to highlight the differ-  
455 ences in estimates of both the group mean and group standard deviations.

456 For both the all-subjects and outliers-removed subject groups, the variance  
457 weighted mean is approximately lower across the artifact above the corpus  
458 callosum and, to a lesser degree, over the left internal and external cap-  
459 sules. For both groups, standard deviation estimates vary more between  
460 regular and weighted averaging, with a lower weighted average across the  
461 white matter artifact, equal values in most of the white matter and higher  
462 estimates near the border with gray matter. Group statistics with a few out-  
463 lier subjects removed give lower averages and lower standard deviations  
464 for both weighted and regular averaging. Removing the outlier subjects  
465 brings the regular and weighted averages closer to each other, with per-  
466 centile differences dropping by at least half.

467 The white matter artifact is most present in the regular average over all sub-  
468 jects (baseline), followed by regular averaging over the reduced group (ar-  
469 tifact reduction method two), then by weighted averaging over all subjects  
470 (artifact reduction method one), and the artifact is least present in weighted  
471 averaging over the reduced group (artifact reduction method three).

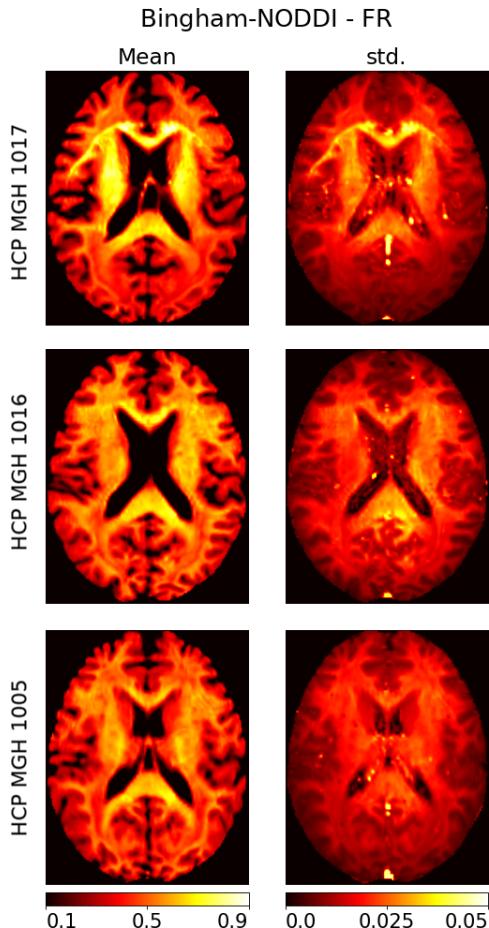


Figure 7: Illustration of artifacts in the HCP MGH datasets using the Bingham-NODDI Fraction of Restricted (FR) mean and standard deviation (std.) estimates from the MLE methodology. In the top row, estimates for HCP MGH subject 1017, with an artifact across the corpus callosum. In the middle row, estimates for HCP MGH subject 1016 with increased standard deviation estimates near a ventricle. In the bottom row, estimates for HCP MGH subject 1016 with no artifacts visible in the mean or standard deviation map.

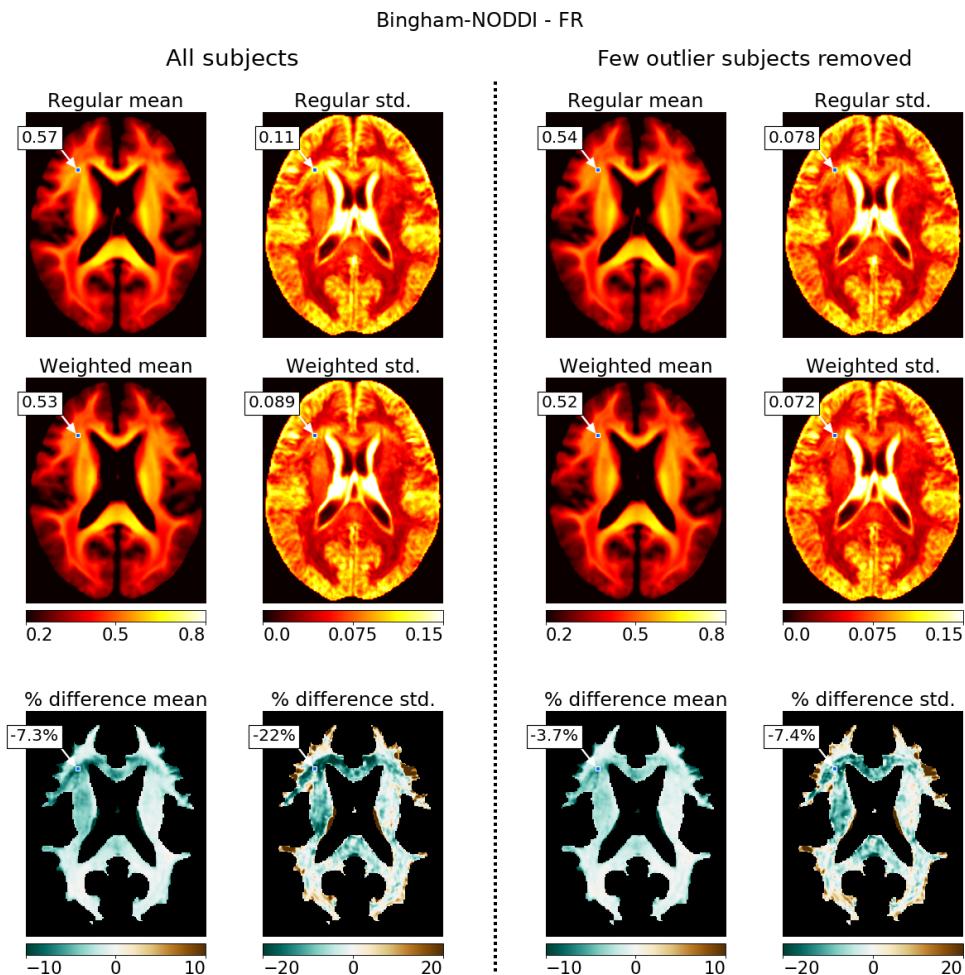


Figure 8: Group averages of Bingham-NODDI Fraction of Restricted (FR) estimates using the HCP MGH data, once over all 35 subjects (left two columns) and once over only 30 subjects where 5 outlier subjects have been removed (right two columns). First row, the regular mean and standard deviation, second row, the variance weighted mean and standard deviations, final row, percentage difference between regular and weighted averages. Point estimates and variances were computed using the MLE methodology.

472 **4 Discussion**

473 We evaluated parameter variance estimates as a quantification of parameter  
474 uncertainties. We compared standard deviation estimates from Maximum  
475 Likelihood Estimation (MLE) plus the Fisher Information Matrix (FIM) to those of Markov Chain Monte Carlo (MCMC) sampling and showed  
476 that both results are identical in ~98.7% of the voxels. In terms of computer  
477 processing time, the estimates of MLE+FIM computed about 38x faster than  
478 those of MCMC. We then showed how data complexity and the signal-to-  
479 noise ratio can affect the parameter variances. Finally, we illustrated how  
480 the parameter variances can be applied in group studies to identify and  
481 downweight the effect of outliers, thereby decreasing the variance in group  
482 estimates, leading to an increase in statistical power of group studies.  
483

484 *4.1 Comparison of the FIM and MCMC*

485 In general, we noted a close correspondence between the parameter dis-  
486 tribution estimates from the FIM and those from MCMC sampling, with  
487 an average similarity of ~98.7% across six models and two datasets. Com-  
488 pared on runtime, computing MLE+FIM is about 38x faster than the use of  
489 MCMC, for comparable results.

490 We made the explicit assumption that the parameter posterior distributions  
491 would follow a Gaussian distribution with a single mode. Theoretically,  
492 only a symmetrical distribution with a single mode would have an equal  
493 mode and mean. Therefore, if the MLE point estimate, which attempts to  
494 find the mode of the posterior, and the MCMC point estimate, which was  
495 computed here as the mean of the sample distribution are equal, then this  
496 is evidence towards symmetric single mode posteriors. Since our results  
497 from the FIM and MCMC were highly comparable (i.e. up to 98.7% of  
498 points estimates were indeed nearly equal), the Gaussian assumption is of-  
499 ten confirmed. In the remaining 1.3% of the voxels, it could either be that  
500 the parameter posteriors were not fully Gaussian distributed, or that the  
501 posterior distributions were multi-modal. In the case of a multi-modal dis-  
502 tribution, the FIM will give variance estimates around a single mode only,  
503 the mode found by the maximum likelihood routine. Our current MCMC  
504 methodology would provide an average and variance over all modes. To  
505 properly deal with multi-modal distributions when using MCMC, would re-  
506 quire fitting a multi-modal normal distribution to the MCMC samples. If  
507 the parameters are not normally distributed, like for example near param-  
508 eter boundaries or with skewed posterior's, the FIM no longer applies and  
509 MCMC would require different post-processing of the samples.

510 Compared on signal-to-noise (SNR), we note that the FIM provides higher  
511 standard deviation estimates at low SNR (< 10) when compared to MCMC.

512 These differences are small and quickly vanish for  $\text{SNR} \geq 10$ . This follows  
513 results from astrophysics, where they recommend a minimum SNR of 10 to  
514 compute variances using the FIM, in gravitational wave assessments (Ro-  
515 driguez et al., 2013).

516 In general, both the FIM and MCMC give comparable answers and both  
517 can be used for computing parameter standard deviations estimates to com-  
518 pute uncertainty. The only essential difference is one of computation time,  
519 computing a maximum likelihood point estimate together with the FIM is  
520 about 38x faster than using MCMC. This was expected, MCMC is generally  
521 known to be a time-consuming process, even when run on a GPU (Harms  
522 & Roebroeck, 2018). MLE on the other hand can be applied very efficiently  
523 using a GPU (Harms et al., 2017) and computing the FIM requires only a  
524 few extra function evaluations (dependent on the number of parameters,  
525 see Appendix A).

#### 526 4.2 *Effects on estimates of the standard deviations*

527 There are several model and data characteristics that can affect standard  
528 deviation estimates, like data complexity, derived parameter maps and the  
529 signal-to-noise ratio. In general, these effects apply equally to both the FIM  
530 and MCMC.

531 Concerning data dependency, as expected, standard deviation estimates  
532 on the RLS-pilot dataset are generally higher than those on the HCP MGH  
533 dataset, reflecting a decrease in point estimate uncertainties with more data  
534 points. The same holds for the relatively large standard deviations in the  
535 Tensor Fractional Anisotropy (FA) estimates, since for the Tensor model we  
536 used only the data volumes with a low b-value.

537 A higher variance can additionally be observed for parameter maps which  
538 are not estimated directly but derived from the estimated parameters. This  
539 makes the variance of such derived parameters maps also a function of  
540 multiple variances, often leading to a higher total variance. This can for ex-  
541 ample be observed in the Tensor FA measure. The same compound effect  
542 could apply to the variance of the Fraction of Stick (FS) of the Ball&Stick  
543 models. For an increasing number of Sticks, the variance in FS is also a  
544 function of multiple volume fractions, which could increase the total vari-  
545 ance.

546 For all models, parameter standard deviations are influenced by the signal-  
547 to-noise (SNR) ratio of the data, with a low SNR ( $< 10$ ) leading to a large  
548 increase in standard deviations. Both shown in real and simulated data, the  
549 effect of SNR on the standard deviation estimates seems to be more gradual  
550 after an  $\text{SNR} \geq 20$ .

551 4.3 *Artifact detection*

552 The computed parameter standard deviations (either from the FIM or MCMC)  
553 could be used as a tool for detecting acquisition artifacts. In one provided  
554 example (figure 7 top row), an artifact in the white matter was visible in  
555 both the parameter estimate and in the standard deviation as a patch of  
556 high intensity voxels. In another example (figure 7 middle row), a patch of  
557 high intensity voxels was visible in the standard deviation estimate but not  
558 in the parameter estimate itself. As such, standard deviation maps have  
559 the potential to be more sensitive in detecting white matter artifacts than  
560 point estimate maps themselves. A promising future development could  
561 be to include these standard deviation maps into quality control frame-  
562 works (Bastiani et al., 2019; Liu et al., 2010; Oguz et al., 2014).

563 4.4 *Increasing power in group studies*

564 By weighing down voxels with a high standard deviation, weighted aver-  
565 aging can reduce the effect of white matter artifacts, approach lower and  
566 more accurate estimates of group variances and increase power of group  
567 statistics. In theory, if the within group datapoints are distributed with  
568 the same mean, variance weighted averaging promises the lowest possi-  
569 ble variance in the group mean. We observe this in large parts of the white  
570 matter where weighted averaging lowers the variance in the group average  
571 as expected, thereby indirectly increasing power in group comparisons.

572 We have shown that some white matter artifacts are visible in the parame-  
573 ter standard deviation maps as patches of relatively large standard devi-  
574 tions. Since variance weighted averaging automatically reduces the effects  
575 of outliers whenever they have a large variance, variance weighted averag-  
576 ing automatically reduces the presence of artifacts. Even after removing a  
577 few subjects with a similar artifact, white matter averaging still reduces the  
578 presence of what appears to be a lower-expressed artifact in the remaining  
579 subjects. Due to this mechanism, subjects no longer need to be excluded  
580 from analysis, thereby improving the power of one's study.

581 Near the gray-white matter border we noticed some voxels where weighted  
582 averaging provides a higher variance than regular averaging. Theoretically,  
583 weighted averaging only predicts lower standard deviations if the points  
584 are distributed with the same mean. Misalignment between subjects can  
585 cause a single voxel to contain white matter for one subject and gray mat-  
586 ter for another subject. Parameter estimates on such voxels will then be  
587 distributed with a different mean, leading to a higher group standard devi-  
588 ation when applying weighted averaging. This could be considered to be  
589 desirable, since such misalignment should not lead to high certainty group  
590 results and is therefore downweighted by the weighted averaging. In other

591 words, the weighted group standard deviation could diagnose alignment  
592 errors in group studies.

593 We note that although weighted averaging is shown here over subjects,  
594 weighted averaging can also be applied within subjects. For example, when  
595 averaging voxels over a white matter tract. In essence, weighted averaging  
596 can be applied in all cases where variances of an estimate are available. In  
597 the future this could be applied to tract based microstructure or tractometry  
598 studies (Bells et al., 2011), for tract based summary statistics with a lower  
599 variance.

## 600 5 Conclusions and recommendations

601 Considering the advantages in processing time and close correspondence  
602 to Markov Chain Monte Carlo estimates, we recommend the use of the  
603 Fisher Information Matrix theory to quantify the uncertainties in parame-  
604 ter estimates. In individual subjects, the parameter standard deviations can  
605 help in detecting white matter artifacts as patches of relatively large stan-  
606 dard deviations. In group statistics, we recommend using the parameter  
607 standard deviations by means of variance weighted averaging. Doing so  
608 can reduce the overall variance in group statistics and reduce the effect of  
609 data artifacts without discarding data from the analysis. Both these effects  
610 can lead to a higher statistical power in group studies.

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815 **Appendix A Numerical Hessian**

816 To compute the Hessian we use a numerical differentiation routine with  
 817 multiple step sizes and extrapolations to provide an estimate with a  $\mathcal{O}(h^6)$   
 818 order of accuracy. For a single step size vector  $\mathbf{d}$ , we compute each element  
 819 of the Hessian using a second order Taylor expansion central difference,

$$\mathbf{H}_{ij}(\mathbf{x}) := \frac{1}{4\mathbf{d}_i\mathbf{d}_j} [ \begin{aligned} & l(\mathbf{x} + \mathbf{e}_i\mathbf{d}_i + \mathbf{e}_j\mathbf{d}_j) \\ & -l(\mathbf{x} + \mathbf{e}_i\mathbf{d}_i - \mathbf{e}_j\mathbf{d}_j) \\ & -l(\mathbf{x} - \mathbf{e}_i\mathbf{d}_i + \mathbf{e}_j\mathbf{d}_j) \\ & +l(\mathbf{x} - \mathbf{e}_i\mathbf{d}_i - \mathbf{e}_j\mathbf{d}_j) \end{aligned} ] \quad (\text{A.1})$$

820 where  $\mathbf{x} \in \mathcal{R}^n$  is the parameter vector,  $l(\mathbf{x})$  is the log-likelihood function  
 821 and  $\mathbf{e}_k$  is a zeros vector with only element  $k$  set to one. We evaluate the  
 822 Hessian multiple times with exponentially diminishing steps and with the  
 823 largest step size chosen such that  $\mathbf{x} \pm \mathbf{d}$  is within bounds and  $\mathbf{d}$  is within  
 824 predefined upper and lower limits. In this work we evaluate the Hessian  
 825 for five different step sizes  $\mathbf{d}$  with each step half the previous step. We  
 826 then apply Richardson extrapolation (Burg & Erwin, 2009) twice to produce  
 827 three estimates with a sixth order of accuracy. These three approximations  
 828 we extrapolate again using Wynn's epsilon algorithm (Weniger, 1991) to  
 829 arrive at a single final estimate.

830 **Appendix B Uncertainty propagation**

831 This appendix provides two illustrations of uncertainty propagation, one  
 832 example using Ball&Stick Fraction of Stick and one example using Tensor  
 833 Fractional Anisotropy.

834 Uncertainty propagation of the Ball&Stick Fraction of Stick can be defined  
 835 as follows. For a two Stick Ball&Stick model, the Fraction of Stick is defined  
 836 as:

$$\text{FS} = w_0 + w_1 \quad (\text{B.1})$$

837 The analytical gradient of this function is given by:

$$\nabla_{\text{FS}} = (w_0, w_1) \quad (\text{B.2})$$

838 The covariance matrix of the weights can be defined as:

$$\Sigma_w = \begin{pmatrix} \sigma_{w_0}^2 & \sigma_{w_0 w_1} \\ \sigma_{w_1 w_0} & \sigma_{w_1}^2 \end{pmatrix} \quad (B.3)$$

839 with  $\sigma_{w_i}^2$  denoting the variance of weight  $w_i$ , and  $\sigma_{w_i w_j}$  denoting the co-  
840 variances of weights  $w_i$  and  $w_j$ . When evaluated, these quantities are taken  
841 from the covariance matrix provided by the FIM.

842 Using equation 6, we can write the uncertainty propagation as:

$$\sigma_{FS}^2 = \nabla_{FS} \Sigma_w \nabla_{FS}^\top \quad (B.4)$$

843 which simplifies to:

$$\sigma_{FS}^2 = w_0^2 \sigma_{w_0}^2 + w_1^2 \sigma_{w_1}^2 + 2w_0 w_1 \sigma_{w_0 w_1} \quad (B.5)$$

844 By evaluating expression B.5 using the point estimates, variance estimates  
845 and covariance estimates of the weights, we can compute the variance in  
846 the FS metric.

847 Uncertainty propagation of Tensor FA is slightly more complex considering  
848 FA is not a linear function of its inputs. The Tensor FA can be defined  
849 in terms of the three Tensor diffusivities (the eigenvalues of the diffusion  
850 Tensor) as:

$$FA = \sqrt{\frac{1}{2} \frac{\sqrt{(d_0 - d_1)^2 + (d_1 - d_2)^2 + (d_0 - d_2)^2}}{\sqrt{d_0^2 + d_1^2 + d_2^2}}} \quad (B.6)$$

851 The derivative of FA with respect to the first diffusivity can be written as:

$$\frac{\partial FA}{\partial d_0} = \frac{2d_0 d_1 d_2 + d_0^2 (d_1 + d_2) - d_1^2 d_2 - d_1 d_2^2 - d_1^3 - d_2^3}{2 \sqrt[3/2]{d_0^2 + d_1^2 + d_2^2} \sqrt{d_0^2 - d_0(d_1 + d_2) + d_1^2 - d_1 d_2 + d_2^2}} \quad (B.7)$$

852 and similar derivatives can be derived for the second and third diffusivity  
853 by suitable permutations of the diffusivity indices. The analytical gradient  
854 of FA,  $\nabla_{FA}$  can now be defined as:

$$\nabla_{FA} = \left( \frac{\partial FA}{\partial d_0}, \frac{\partial FA}{\partial d_1}, \frac{\partial FA}{\partial d_2} \right) \quad (B.8)$$

855 The covariance matrix of the diffusivities can be defined as:

$$\Sigma_d = \begin{pmatrix} \sigma_{d_0}^2 & \sigma_{d_0d_1} & \sigma_{d_0d_2} \\ \sigma_{d_1d_0} & \sigma_{d_1}^2 & \sigma_{d_1d_2} \\ \sigma_{d_2d_0} & \sigma_{d_2d_1} & \sigma_{d_2}^2 \end{pmatrix} \quad (B.9)$$

856 with  $\sigma_{d_i}^2$  denoting the variance of diffusivity  $d_i$ , and  $\sigma_{d_i d_j}$  denoting the co-  
857 variances of diffusivities  $d_i$  and  $d_j$ .

858 Using equation 6, we can define the uncertainty propagation of FA as:

$$\sigma_{FA}^2 = \nabla_{FA} \Sigma_d \nabla_{FA}^\top \quad (B.10)$$

859 By evaluating expression B.10 using the point estimates of the diffusivities  
860 together with the corresponding variance and covariance estimates from  
861 the FIM, we can compute the propagated variance in the FA metric.